

Search or Article-id (Help | Advanced search) arXiv.org > astro-ph > arXiv:1107.1346 All papers Go! Ŧ Astrophysics > High Energy Astrophysical Phenomena Download: PDF **Are low luminosity GRBs** PostScript Other formats generated by relativistic jets? Current browse context: astro-ph.HE Omer Bromberg, Ehud Nakar, Tsvi Piran < prev | next > new | recent | 1107 (Submitted on 7 Jul 2011) Change to browse by: Low luminosity gamma-ray bursts (II-GRBs) constitute a sub-class of gammaastro-ph ray bursts (GRBs) that plays a central role in the GRB-supernova connection. While II-GRBs differ from typical long GRBs (LGRBs) in many aspects, they References & Citations also share some common features. Therefore, the question whether the **INSPIRE HEP** gamma-ray emission of II-GRBs and LGRBs has a common origin is of great (refers to | cited by) interest. Here we address this question by testing whether II-GRBs, like NASA ADS LGRBs according to the Collapsar model, can be generated by relativistic jets that punch holes in the envelopes of their progenitor stars. The collapsar Bookmark(what is this?) model predicts that the durations of most observed bursts will be comparable 📃 💿 🗶 💀 🖬 🔚 📲 🔛 🧐 to, or longer than, the time it takes the jets to breakout of the star. We calculate the jet breakout times of II-GRBs and compare them to the observed durations. We find that there is a significant access of II-GRBs with durations that are much shorter than the jet breakout time and that these are inconsistent with the Collapsar model. We conclude that the processes that dominate the gamma-ray emission of II-GRBs and of LGRBs are most likely fundamentally different.

Comments:Submitted to ApJ letters. 12 pages, 1 figureSubjects:High Energy Astrophysical Phenomena (astro-ph.HE)Cite as:arXiv:1107.1346 [astro-ph.HE](or arXiv:1107.1346v1 [astro-ph.HE] for this version)

## **Submission history**

From: Omer Bromberg [view email] [v1] Thu, 7 Jul 2011 11:18:27 GMT (17kb)

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