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Yoshiaki Ono, Masami Ouchi, Bahram Mobasher, Mark Dickinson, Kyle Penner, Kazuhiro Shimasaku, Benjamin J. Weiner, Jeyhan S. Kartaltepe, Kimihiko Nakajima, Hooshang Nayyeri, Daniel Stern, Nobunari Kashikawa, Hyron Spinrad

(Submitted on 15 Jul 2011 (v1), last revised 21 Nov 2011 (this version, v2))

We present the results of our ultra-deep Keck/DEIMOS spectroscopy of zdropout galaxies in the SDF and GOODS-N. For 3 out of 11 objects, we detect an emission line at ~ 1um with a signal-to-noise ratio of ~ 10. The lines show asymmetric profiles with high weighted skewness values, consistent with being Lya, yielding redshifts of z=7.213, 6.965, and 6.844. Specifically, we confirm the z=7.213 object in two independent DEIMOS runs with different spectroscopic configurations. The z=6.965 object is a known Lya emitter, IOK-1, for which our improved spectrum at a higher resolution yields a robust skewness measurement. The three z-dropouts have Lya fluxes of 3 x 10^-17 erg s^-1 cm^-2 and rest-frame equivalent widths  $EW_0^{-1}$  = 33-43A. Based on the largest spectroscopic sample of 43 z-dropouts that is the combination of our and previous data, we find that the fraction of Lya-emitting galaxies (EW\_0^Lya > 25A) is low at z ~ 7; 17 +- 10% and 24 +- 12% for bright (Muv ~= -21) and faint (Muv ~= -19.5) galaxies, respectively. The fractions of Lyaemitting galaxies drop from  $z \sim 6$  to 7 and the amplitude of the drop is larger for faint galaxies than for bright galaxies. These two pieces of evidence would indicate that the neutral hydrogen fraction of the IGM increases from z ~ 6 to 7, and that the reionization proceeds from high- to low-density environments, as suggested by an inside-out reionization model.

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Astrophysics > Cosmology and Extragalactic Astrophysics

## **Spectroscopic Confirmation of** Three z-Dropout Galaxies at z = 6.844 - 7.213: Demographics of Lyman-Alpha Emission in z ~ 7 Galaxies

#### Submission history

From: Yoshiaki Ono [view email] [v1] Fri, 15 Jul 2011 20:19:33 GMT (250kb) [v2] Mon, 21 Nov 2011 07:55:10 GMT (270kb)

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