



Emission Measure Distribution and Heating of Two Active Region Cores

Durgesh Tripathi, James A. Klimchuk, Helen E. Mason

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Using data from the Extreme-ultraviolet Imaging Spectrometer aboard Hinode, we have studied the coronal plasma in the core of two active regions. Concentrating on the area between opposite polarity moss, we found emission measure distributions having an approximate power-law form $EM \propto T^{2.4}$ from $\log T = 5.5$ up to a peak at $\log T = 6.55$. We show that the observations compare very favorably with a simple model of nanoflare-heated loop strands. They also appear to be consistent with more sophisticated nanoflare models. However, in the absence of additional constraints, steady heating is also a viable explanation.

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