

896A

Cornell University Library

arXiv.org > astro-ph > arXiv:1107.3180

Astrophysics > Solar and Stellar Astrophysics

Radio Interferometric Planet

Search II: Constraints on sub-

Jupiter-Mass Companions to GJ

Search or Article-id

All papers 🚽 Go!

(Help | Advanced search)

Download:

- PDF
- PostScript
- Other formats

Current browse context: astro-ph.SR

< prev | next >

new | recent | 1107

Change to browse by:

astro-ph astro-ph.EP

References & Citations

- INSPIRE HEP (refers to | cited by)
- NASA ADS

Bookmark(what is this?) E 😳 🎇 🐼 💽 📊 📲 🔛 🗐

Geoffrey C. Bower, Alberto Bolatto, Eric B. Ford, Adam Fries, Paul Kalas, Karol Sanchez, Phoebe Sanderbeck, Vincent Viscomi

(Submitted on 15 Jul 2011)

We present results from the Radio Interferometric Planet (RIPL) search for compan- ions to the nearby star GJ 896A. We present 11 observations over 4.9 years. Fitting astrometric parameters to the data reveals a residual with peak-to-peak amplitude of ~ 3 mas in right ascension. This residual is well-fit by an acceleration term of 0.458 \pm 0.032 mas/y^2. The parallax is fit to an accuracy of 0.2 mas and the proper motion terms are fit to accuracies of 0.01 mas/y. After fitting astrometric and acceleration terms residuals are 0.26 mas in each coordinate, demonstrating that stellar jitter does not limit the ability to carry out radio astrometric planet detection and characterization. The acceleration term originates in part from the companion GJ 896B but the amplitude of the acceleration in declination is not accurately predicted by the orbital model. The acceleration sets a mass upper limit of 0.15 MJ at a semimajor axis of 2 AU for a planetary companion to GJ 896A. For semi-major axes between 0.3 and 2 AU upper limits are determined by the maximum angular separation; the upper limits scale from the minimum value in proportion to the inverse of the radius. Upper limits at larger radii are set by the acceleration and scale as the radius squared. An improved solution for the stellar binary system could improve the exoplanet mass sensitivity by an order of magnitude.

Comments:Accepted for publication in ApJSubjects:Solar and Stellar Astrophysics (astro-ph.SR); Earth and
Planetary Astrophysics (astro-ph.EP)Cite as:arXiv:1107.3180 [astro-ph.SR]

(or arXiv:1107.3180v1 [astro-ph.SR] for this version)

Submission history

From: Geoffrey C. Bower [view email] [v1] Fri, 15 Jul 2011 23:18:32 GMT (46kb) Link back to: arXiv, form interface, contact.